Comment on "The Cosmic Time in Terms of the Redshift", by Carmeli et al.

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Abstract

The time-redshift relation of Carmeli et al. differs from that of the standard flat Λ CDM model by more than 500 million years for $1 \le z \le 4.5$.

In a recent paper in this journal, Carmeli et al. investigated the relationship between the time t that light from a distant object was emitted and its redshift z [1]. They "start by assuming that the universe is empty of gravitation" and derive the time-redshift relation in this case:

$$t(z) = \frac{2t_0}{1 + (1+z)^2} \,, (1)$$

where t_0 is the age of the universe. (Carmeli et al. denote the age by τ .)

Then they modify Eq. (1) to include gravity. They "adopt the method used in classical general relativity theory" and replace t_0 by H_0^{-1} :

$$t(z) = \frac{2H_0^{-1}}{1 + (1+z)^2},$$
 (2)

because it "seems to be the proper thing to do".

Carmeli et al. write that Eq. (2) is "valid for all redshift values" and "It is hoped that the formula will be useful for identifying objects at the early Universe." They nowhere restrict their equation to specific general relativistic cosmological models.

The purpose of this comment is to point out that Eq. (2) is not a good approximation to the time-redshift relation for the currently favored standard flat Λ CDM cosmological model. Specifically, they differ by over half a billion years for $1 \le z \le 4.5$.

We can see immediately that Eq. (2) is not an exact time-redshift relation. For it does not give $t(0) = t_0$, which is obviously true for every cosmological model.

Grøn [2] has given a closed form expression for the expansion factor S(t) in the standard flat ΛCDM model:

$$S^{3}(t) = \frac{\Omega_{M}}{\Omega_{\Lambda}} \sinh^{2}\left(\frac{3H_{0}\Omega_{\Lambda}^{1/2}}{2}t\right) S^{3}(t_{o}), \tag{3}$$

where H_0 is Hubble's constant, Ω_M is the fractional matter energy density today, and Ω_{Λ} is the fractional dark energy density today.

Combining this with $z + 1 = S(t_o)/S(t)$ gives

$$t(z) = \frac{2H_0^{-1}}{3\Omega_{\Lambda}^{1/2}} \sinh^{-1} \left[\left(\frac{\Omega_{\Lambda}}{\Omega_M} \right)^{1/2} (z+1)^{-3/2} \right].$$
 (4)

The currently accepted best values of the parameters in Eq. (4) are $H_0 = 72$ km/sec/Mpc, $\Omega_M = .26$, and $\Omega_{\Lambda} = .74$ [3]. (There is no consensus about the nature of dark matter and dark energy. But there is a wide consensus about the values of Ω_M and Ω_{Λ} .) Substituting these values shows that the difference between Eqs. (2) and (4) is over half a billion years for $1 \le z \le 4.5$.

References

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- [3] D. Spergel, et al., "Wilkinson Microwave Anisotropy Probe (WMAP) Three Year Results: Implications for Cosmology", ApJ submitted, astro-ph/0603449.